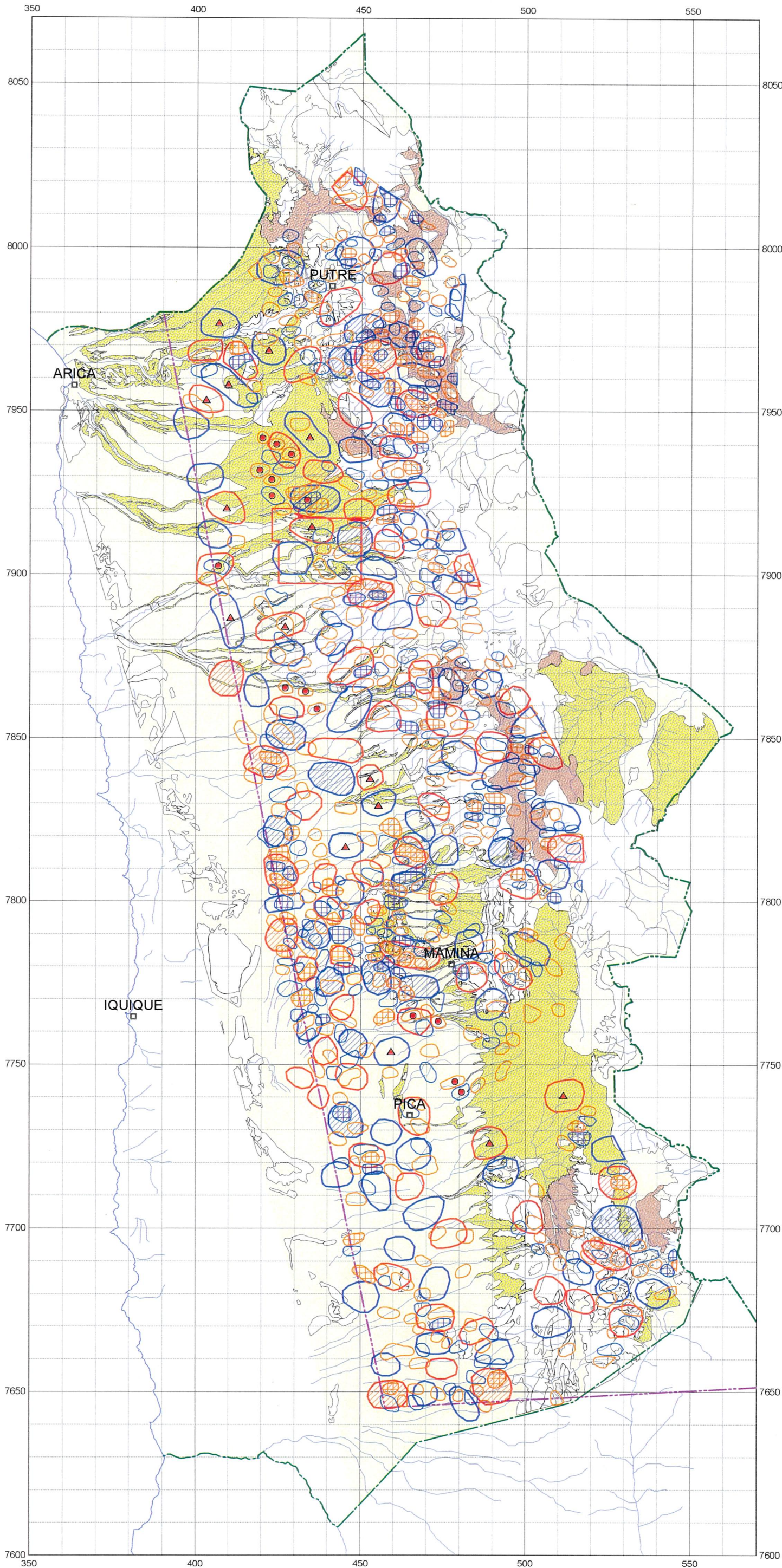


LEGEND

- Short wavelength High
- Short wavelength Low
- Medium wavelength High
- Medium wavelength Low
- [Fill pattern : Large amplitude anomaly]
- ▨ Quaternary volcanics (Qv)
- Anomaly corresponding to Qv and Topography
- ▲ Anomaly corresponding to Topography

Fig.2-3-26
SW&MW Magnetic Anomalies and
Quaternary Volcanics





LEGEND

- Short wavelength High
- Short wavelength Low
- Medium wavelength High
- Medium wavelength Low

[Fill pattern : Large amplitude anomaly]

- Quaternary sediments
- Quaternary ignimbrite
- Tertiary ignimbrite (Tig)
- Anomaly corresponding to Tig and Topography
- ▲ Anomaly estimated to be due to Tig

Fig.2-3-27
SW&MW Magnetic Anomalies and
Tertiary-Quaternary Ignimbrites



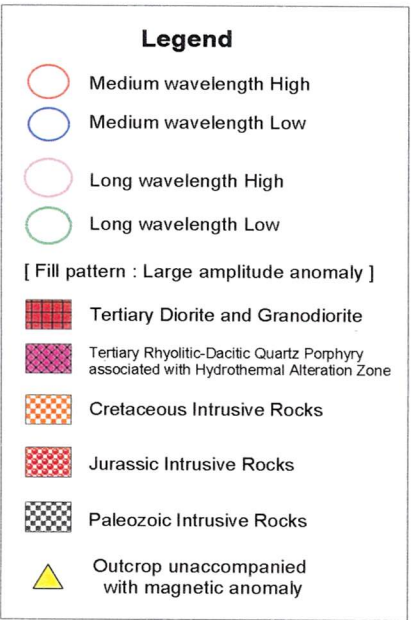
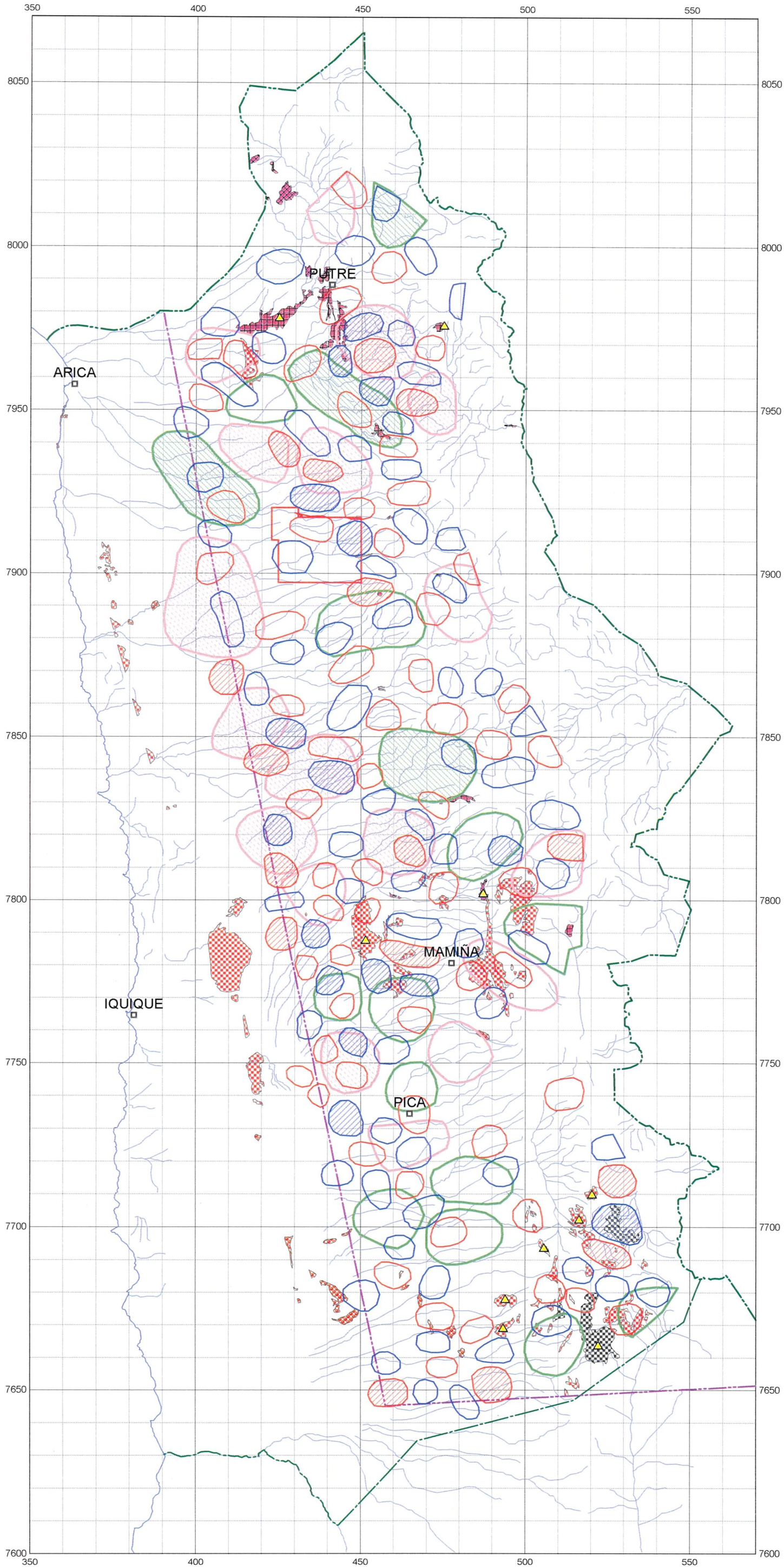
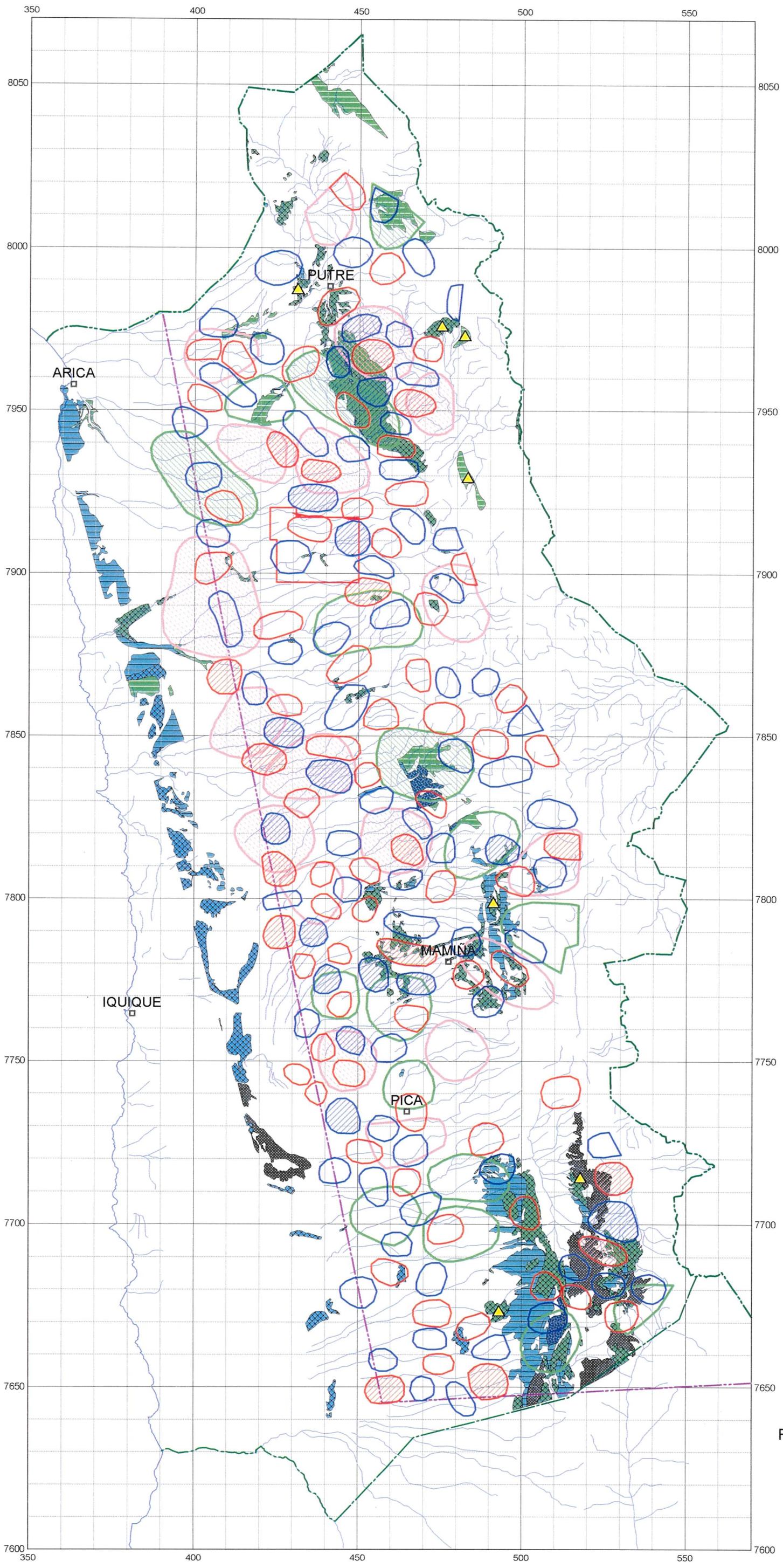


Fig.2-3-28
MW&LW Magnetic Anomalies
and Intrusive Rocks





LEGEND

- Medium wavelength High
- Medium wavelength Low
- Long wavelength High
- Long wavelength Low

[Fill pattern : Large amplitude anomaly]

- Cretaceous - Tertiary sedimentary rocks
- Cretaceous - Tertiary volcanic rocks
- Jurassic sedimentary rocks
- Jurassic volcanic rocks
- Paleozoic sedimentary rocks
- Precambrian rocks
- ▲ Outcrop unaccompanied with magnetic anomaly

Fig.2-3-29
MW&LW Magnetic Anomalies and
Basement Rocks



3-5 Two-Dimensional Modeling

Two dimensional modeling regarding magnetic susceptibility distribution was carried out along 30km long lines extending through 12 drilling sites. The drilling was done in the northern half of the survey area. The drilling sites and lines for modeling were plotted on the total magnetic intensity map, and it is shown in Figure 2-3-30. Also drilling sites and modeling lines were plotted on a geologic map and on a reduced to the pole magnetic map, and they are shown in Figures 2-3-31 and 2-3-32. The modeling lines were set in the N-S direction in line with the direction of the magnetic field.

Modeling was carried out by dividing the subsurface area into three layers. Short wavelength total magnetic intensity was used to be fitted for the first layer, medium wavelength total magnetic intensity for the second layer, and long wavelength total magnetic intensity for the third layer. The upper surface of the first layer was placed 300m below the earth's surface, while the lower surface was placed 1,800m below the assumed flight height (= calculation point). The upper surface of the second layer was set to coincide with the lower surface of the first layer and the lower surface was placed at a constant depth so that the average thickness of this layer would be about 2,500m. The upper surface of the third layer was set to coincide with the lower surface of the second layer and the lower surface was placed at 30km below sea level. The thickness of the above layers were determined, after some trials and errors, to satisfy conditions such as: to be within harmonious range of the magnetic susceptibility values obtained by modeling and those actually measured, and extreme gap of magnetic susceptibility does not exist between adjacent layers.

Each layer was divided into blocks of: 1km width for the first layer, 2km width for the second, and 3km for the third. The magnetic susceptibility of each block was assumed to be homogeneous. For the second and third layers, calculated and measured magnetic anomaly values were fit by changing the magnetic susceptibility but without changing the shape of the blocks. But for the first layer, improvement of fitting was attempted by changing the shape of the upper and side surfaces as well as the magnetic susceptibility. For the zone between the earth's surface and the first layer, magnetic susceptibility value of $1.0\sim 5.0 \times 10^{-3}$ was assumed from the drilling data.

The results of the 2-dimensional modeling is shown in Figures 2-3-33 to 2-3-44. The calculated magnetic susceptibility values are shown in these figures and these values are grouped into 8 ranks which are shown in different colors. Also assumed structural lines are drawn for zones with remarkable horizontal gap of magnetic susceptibility, and boundary lines are drawn for parts where layered or massive structure can be distinguished in susceptibility distribution to clarify their existence.

Regarding Lines 3, 4, 5, and 6, a confluence of blocks with high susceptibility is largely raised from the third to the second layer, and the shape of the rise indicates the possibility of the high magnetic body being an intrusive rock body.

Regarding Lines 7, 8, and 10; these lines contain strong magnetic blocks in some parts. But on the whole, a confluence of not very strong ($7.5\sim 10\times 10^{-3}$) magnetic blocks is raised from the third to the second layer. Magnetism of this confluence is not strong, but the shape indicates the possibility of an intrusive body.

Lines 1, 11, 12 are lines where basement rocks have been confirmed by drilling. The following common structure occurs along these lines. A confluence of blocks with weak negative susceptibility (reverse remanent magnetism) occurs in the third or third to second layer, and this is surrounded by not very strong positive blocks. The meaning of such magnetic susceptibility distribution is not clear at this point, but the shape indicates the possibility of an intrusive body.

The confluences of blocks with laterally continuous magnetic susceptibility along the first layer or from the first to the second layer were obtained in many modeling lines. Such confluences are considered to be, from the form of distribution, very possibly expressions of ignimbrite occurring on the surface. Some confluences inferred to be ignimbrite show high magnetic susceptibility. This generally appear to be incompatible to the magnetism of ignimbrites. But results of remanent magnetism measurements show that they can cause the strong magnetic anomalies similar to those by intrusive rocks. Thus the above is not necessarily contradictory.

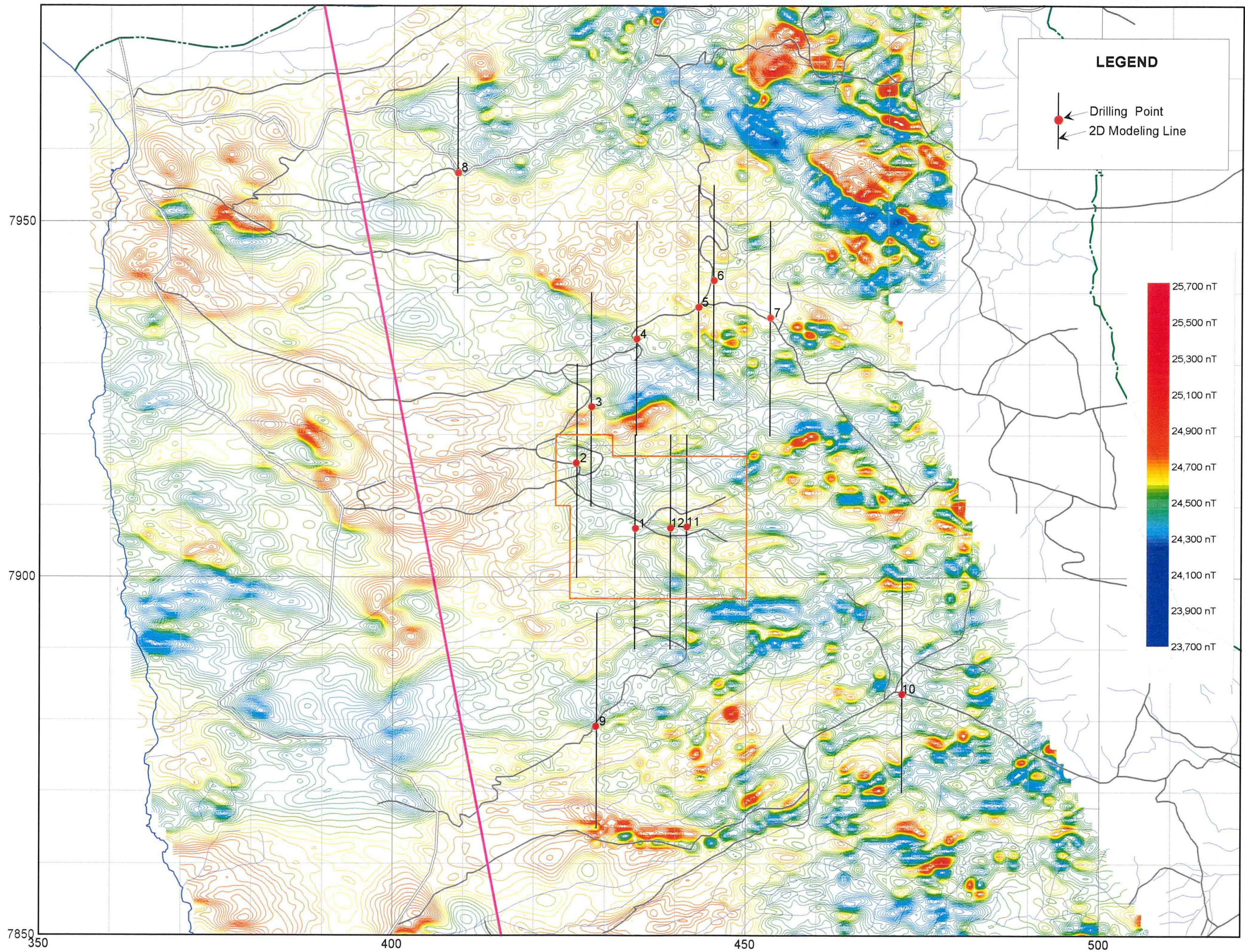


Fig.2-3-30 Location of 2D Modelling Lines and Total Magnetic Intensity

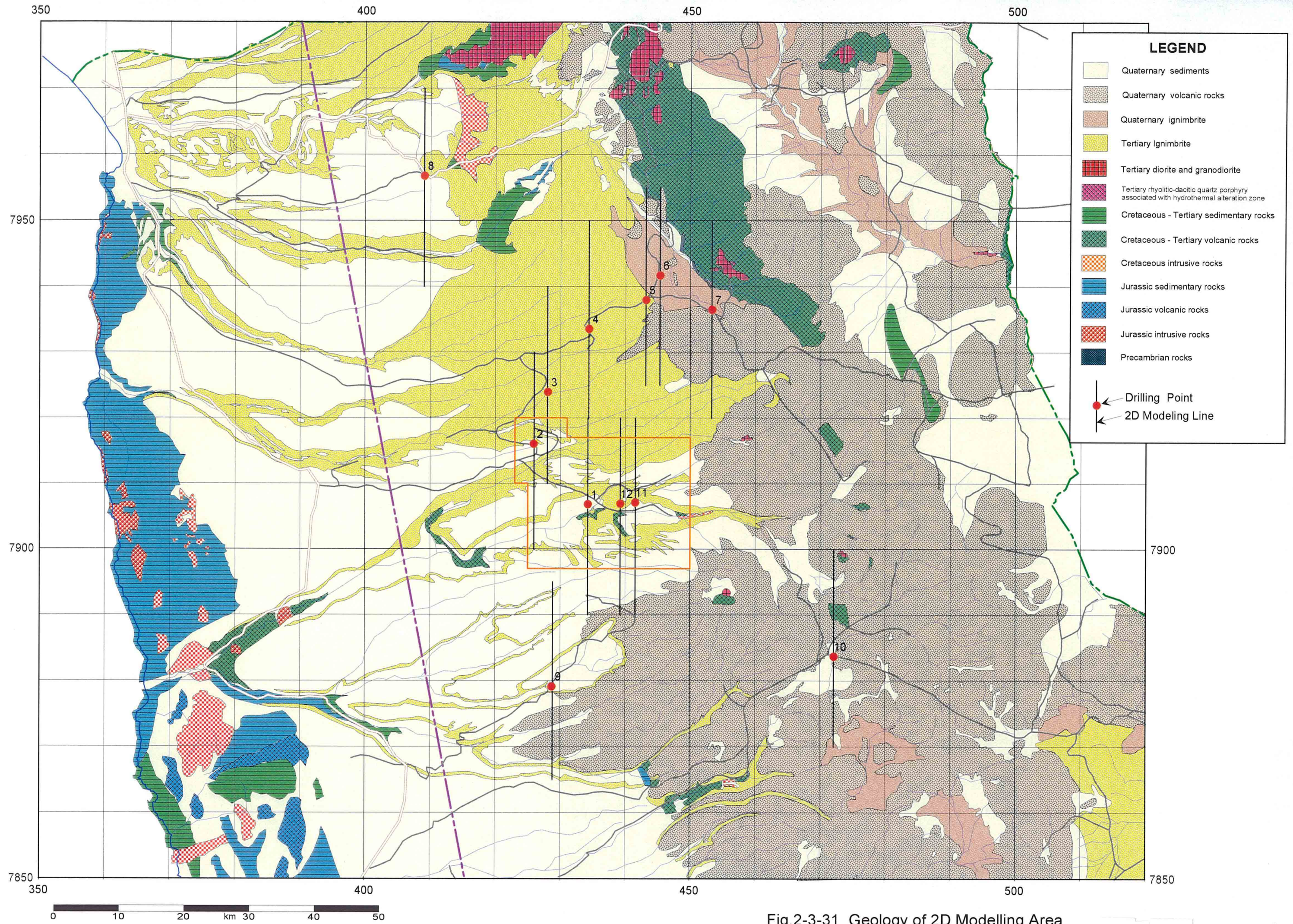


Fig.2-3-31 Geology of 2D Modelling Area

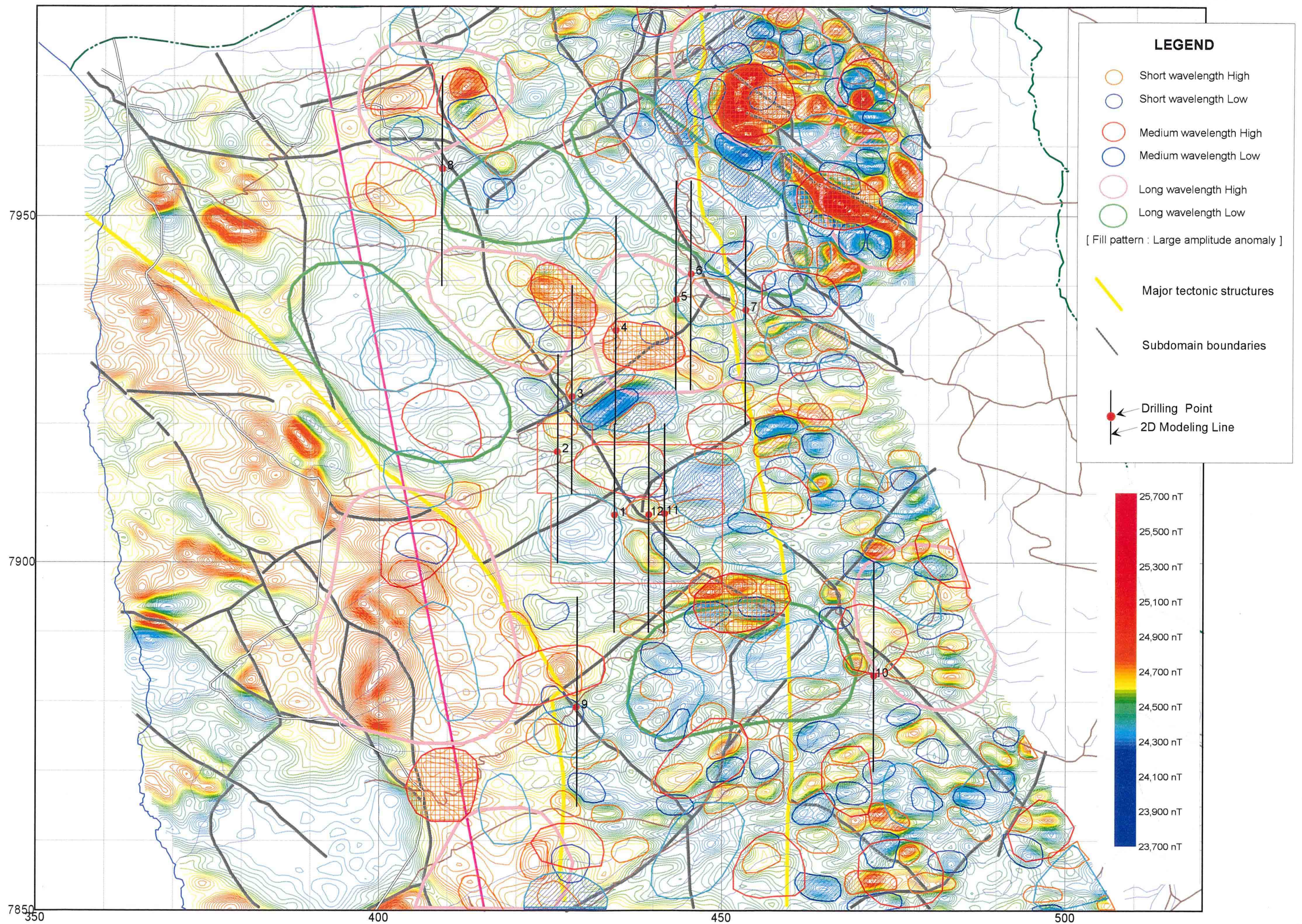


Fig.2-3-32 Reduced to the Pole and Magnetic Structures of 2D Modelling Area

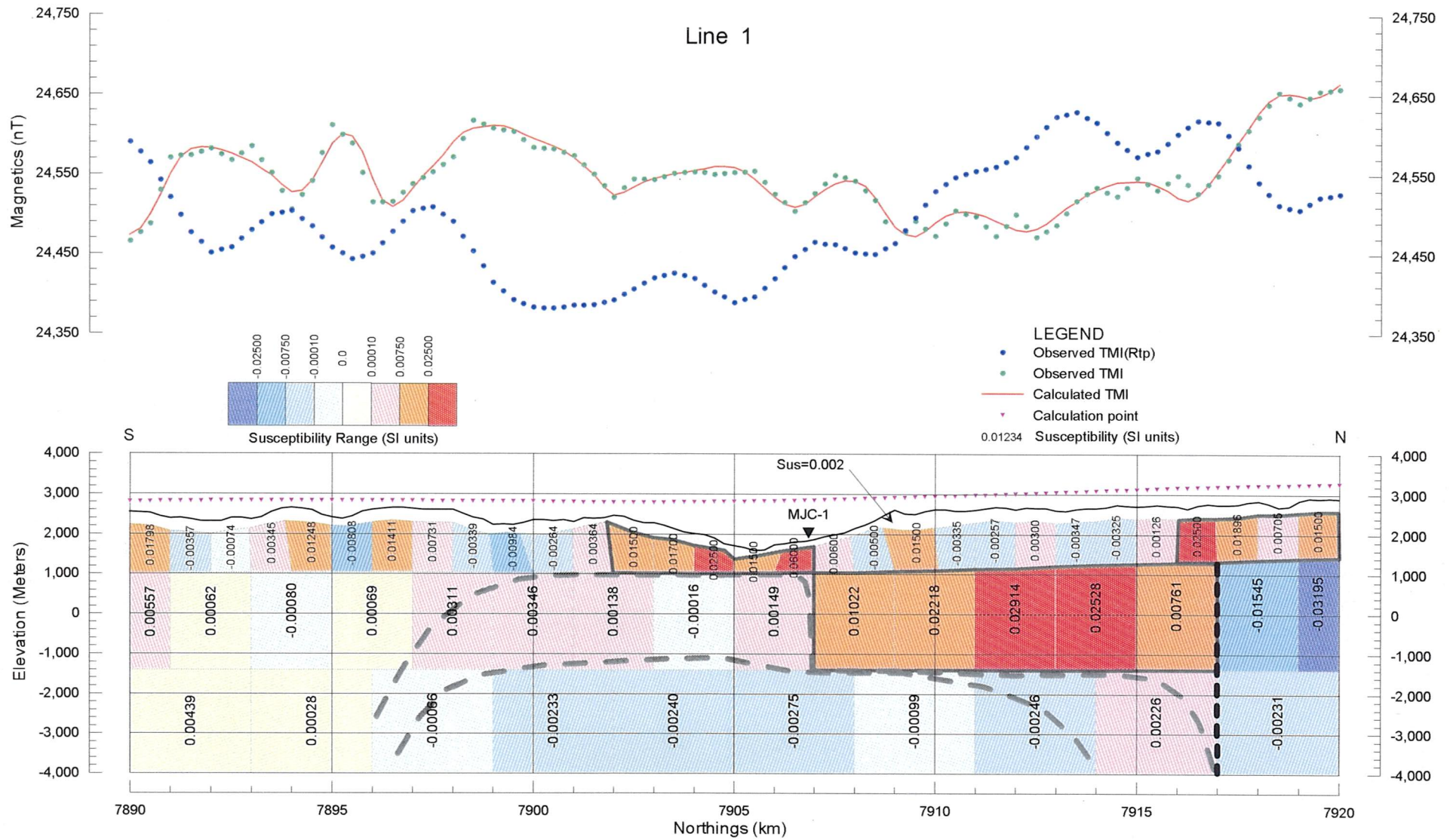


Fig. 2-3-33 Line 1 - 2D Model Profile

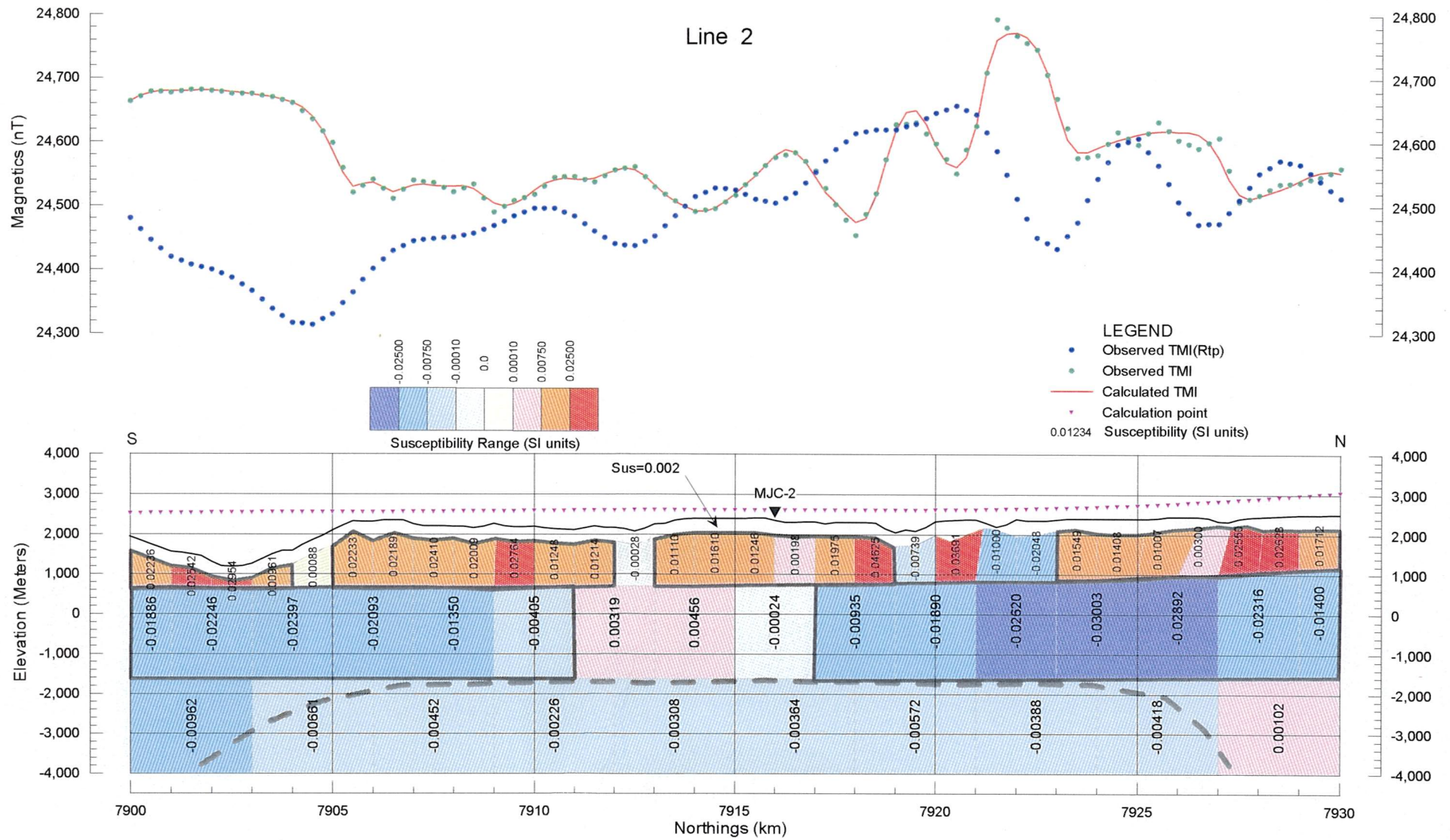


Fig. 2-3-34 Line 2 - 2D Model Profile

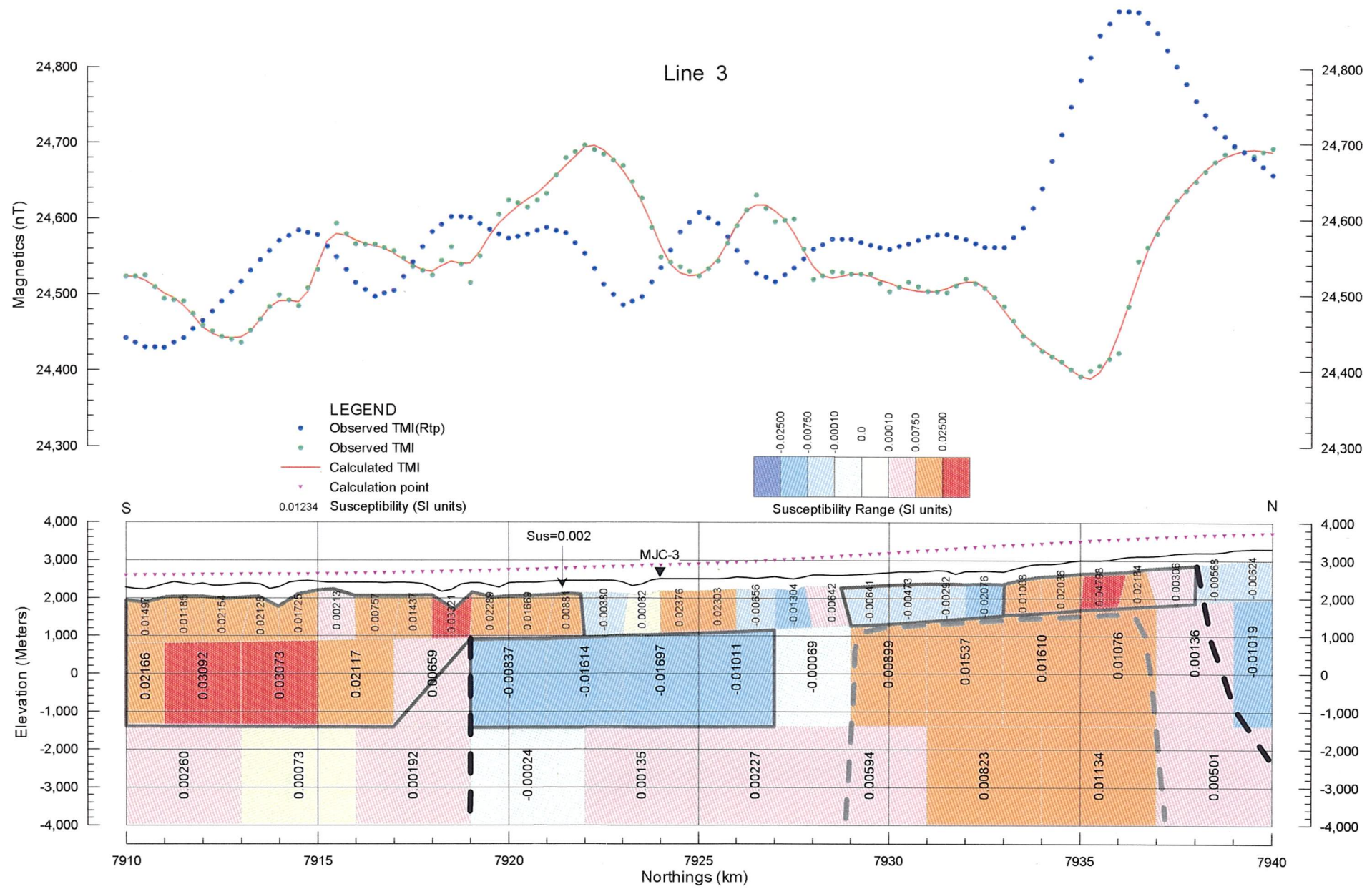


Fig. 2-3-35 Line 3 - 2D Model Profile